Scaling exponents in weakly anisotropic turbulence from the Navier–Stokes equation

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(Received 16 February 2001 and in revised form 13 April 2001)

The second-order velocity structure tensor of weakly anisotropic strong turbulence is decomposed into its SO(3) invariant amplitudes $d_j(r)$. Their scaling is derived within a scaling approximation of a variable-scale mean-field theory of the Navier–Stokes equation. In the isotropic sector j = 0 Kolmogorov scaling $d_0(r) \propto r^{2/3}$ is recovered. The scaling of the higher j amplitudes (j even) depends on the type of the external forcing that maintains the turbulent flow. We consider two options: (i) for an analytic forcing and for decreasing energy input into the sectors with increasing j, the scaling of the higher sectors j > 0 can become as steep as $d_j(r) \propto r^{j+2/3}$; (ii) for a non-analytic forcing we obtain $d_j(r) \propto r^{4/3}$ for all non-zero and even j.

1. Introduction

In the last few decades scaling in fully developed turbulence has mainly been analysed in terms of the longitudinal velocity structure functions (Monin & Yaglom 1975; Frisch 1995). Meanwhile experimental and numerical evidence has accumulated that, at least for finite Reynolds numbers, the transversal structure functions scale differently (Noullez *et al.* 1997; Grossmann, Lohse & Reeh 1997; Dhruva, Tsuji & Sreenivasan 1997; Chen *et al.* 1997; van de Water & Herweijer 1999). Two questions immediately arise: (i) what is the proper decomposition of the velocity structure tensor into invariant amplitudes, and (ii) what is the origin of their different scalings?

In addressing the first question, Arad, L'vov & Procaccia (1999b) suggested decomposing the second-order velocity structure tensor into the amplitudes $d_{jmq}(r)$ of the irreducible SO(3) representation,

$$D_{ik}(\mathbf{r}) = \langle\!\langle v_i(\mathbf{r}, t) v_k(\mathbf{r}, t) \rangle\!\rangle = \sum_{jmq} d_{jmq}(r) B_{ik}^{jmq}(\hat{\mathbf{r}}), \qquad (1.1)$$

reflecting the rotational symmetry of the Navier–Stokes equation. Here, $v_i(\mathbf{r}, t) = u_i(\mathbf{x} + \mathbf{r}, t) - u_i(\mathbf{x}, t)$ is the velocity difference, the brackets $\langle \langle \dots \rangle \rangle$ denote the ensemble average and, as in Arad *et al.* (1999*b*), the tensors $B_{ik}^{jmq}(\hat{\mathbf{r}})$ are combinations of the spherical harmonics $Y_{jm}(\hat{\mathbf{r}})$ and operations like ∂_{r_i} , r_k , δ_{ik} , the index *q* labels the different types of such combinations, and $\hat{\mathbf{r}}$ denotes the unit vector in the direction of $\mathbf{r}, \hat{\mathbf{r}} = \mathbf{r}/r$. Non-zero values of *j* contribute to D_{ik} if the turbulence is not isotropic.

In Arad *et al.* (1998), Arad *et al.* (1999*a*), Kurien & Sreenivasan (2000), Kurien *et al.* (2000), and Kurien & Sreenivasan (2001) the scaling exponents of the amplitudes $d_j(r)$ were extracted from experimental as well as numerical data. For j = 0 a scaling

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exponent close to the Kolmogorov value 2/3 was recovered, but for j = 2 values close to 4/3 were found. This scaling exponent corresponds to a power spectrum $\sim k^{-7/3}$, a behaviour first suggested for shear flow by Lumley (1967) through a dimensional argument. Experimental evidence for it was found by Wyngaard & Cote (1972) and, later, also by Saddoughi & Veeravalli (1994), among others. For higher j > 2 Biferale & Toschi (2001) have found even larger scaling exponents from the analysis of numerical data, namely 1.67–1.7 for j = 4 and 3.3–3.4 for j = 6.

Here we aim at analytically *calculating* the mean field part of the scaling exponents of the *j*-amplitudes from the Navier–Stokes equation for weakly anisotropic, homogeneous turbulence. We employ the variable-scale mean-field theory of Effinger & Grossmann (1987), i.e. we disregard intermittency corrections.

2. The Effinger–Grossmann mean field theory for the weakly anisotropic case

The main idea in Effinger & Grossmann (1987) is to decompose the velocity field into a smooth part $u_i^{(r)}$, defined as spatial average over a sphere with variable radius r and therefore containing only scales larger than r, and a strongly varying part $\tilde{u}_i^{(r)}$, to which the scales smaller than r contribute. Within the Effinger–Grossmann theory, not only can the K41 r-scaling exponent 2/3 of the structure function be analytically calculated from the Navier–Stokes equation, but also the Kolmogorov constant b = 6.3. As we now assume (weak) anisotropy of the flow, we introduce an average which reflects its scale r and, in addition, the direction of the averaging. Therefore, for each component, we choose an average over a line in the \hat{r} -direction with the length 2r,

$$u_i^{(r)}(\boldsymbol{x},t) = \frac{1}{2r} \int_{-r}^{r} u_i(\boldsymbol{x} + y\hat{\boldsymbol{r}}, t) \mathrm{d}y \equiv \langle u_i(\boldsymbol{x} + y\hat{\boldsymbol{r}}, t) \rangle_y^{(r)}.$$
 (2.1)

Correspondingly, $\tilde{u}_i^{(r)}(\mathbf{x}, t) = u_i(\mathbf{x}, t) - u_i^{(r)}(\mathbf{x}, t)$. The upper index \mathbf{r} denotes that these averages not only depend on the scale r, but also on the direction $\hat{\mathbf{r}}$ of averaging, thus on the full vector \mathbf{r} . The lower index y indicates the averaged variable. As in the original spherical averaging case there is a close relation between the second order moments of $u_i^{(r)}$ and the structure tensor $D_{ik}(\mathbf{r})$: $\langle\!\langle u_i^{(r)}u_k^{(r)}\rangle\!\rangle = \langle\!\langle u_i u_k\rangle\!\rangle - \frac{1}{2}\langle\langle D_{ik}(\mathbf{y}_1 + \mathbf{y}_2)\rangle_{y_1}^{(r)}\rangle_{y_2}^{(r)}$. This relation is crucial for the method. For simplicity we use the abbreviation $\mathbf{y} = \mathbf{y}\hat{\mathbf{r}}$. In the above double average y_1 is thus parallel to y_2 .

Eliminating the pressure p gives a non-local term involving the Green function G(x). Inserting the velocity decomposition into the Navier–Stokes equation and averaging, we obtain an equation of motion for the large-scale ('superscale') velocity:

$$\begin{aligned} \partial_{t}u_{i}^{(r)}(\mathbf{x},t) &= -u_{j}^{(r)}(\mathbf{x},t)\partial_{x_{j}}u_{i}^{(r)}(\mathbf{x},t) - \langle \tilde{u}_{j}^{(r)}(\mathbf{x}+y\hat{\mathbf{r}})\partial_{x_{j}}\tilde{u}_{i}^{(r)}(\mathbf{x}+y\hat{\mathbf{r}})\rangle_{y}^{(r)} \\ &+ v\Delta_{x}u_{i}^{(r)}(\mathbf{x},t) + f_{i}^{(r)}(\mathbf{x},t) \\ &+ \int \mathrm{d}^{3}x'G(x')\partial_{x_{i}'}\{u_{k|l}^{(r)}(\mathbf{x}+\mathbf{x}',t)u_{l|k}^{(r)}(\mathbf{x}+\mathbf{x}',t) \\ &+ \langle \tilde{u}_{k|l}^{(r)}(\mathbf{x}+\mathbf{x}'+y\hat{\mathbf{r}},t)\tilde{u}_{l|k}^{(r)}(\mathbf{x}+\mathbf{x}'+y\hat{\mathbf{r}},t)\rangle_{y}^{(r)} \}. \end{aligned}$$
(2.2)

We use the abbreviation $u_{i|k}(\mathbf{x}, t) := \partial_{x_k} u_i(\mathbf{x}, t)$, etc.; Δ_x denotes the Laplacian with respect to \mathbf{x} , \mathbf{v} is the kinematic viscosity, and f_i an external forcing maintaining the turbulent flow. In Effinger & Grossmann (1987) isotropic forcing is considered. Here, by proper choice of f_i we explicitly introduce anisotropy. It implies an anisotropic

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energy input whose characteristic details will be discussed later. Subtracting (2.2) from the Navier–Stokes equation gives an equation for the 'subscale' velocity $\tilde{u}_i^{(r)}$. Its formal solution can be found by time integrating along a Lagrangian path x(t';z,t) of a fluid particle which at time t' = t is at the position x = z. Also, from equation (2.2) we can derive an energy balance equation for the superscales.

The central approximation of the mean-field theory of Effinger & Grossmann (1987) is that the small-scale flow is statistically independent of the smooth large-scale one. Therefore, in higher-order moments we factorize the $u^{(r)}$ from the $\tilde{u}^{(r)}$, e.g.

$$\langle\!\langle \widetilde{u}^{(r)}\widetilde{u}^{(r)}u^{(r)}u^{(r)}\rangle\!\rangle \simeq \langle\!\langle \widetilde{u}^{(r)}\widetilde{u}^{(r)}\rangle\!\rangle\langle\!\langle u^{(r)}u^{(r)}\rangle\!\rangle.$$

Physically this means that the large scales feel the small ones as a kind of eddy viscosity. Note again that this factorization excludes intermittency effects. Another assumption is that in the time integration along a Lagrangian path of a fluid particle the slow t'-dependence of the superscales $u^{(r)}$ is neglected since the subscales $\tilde{u}^{(r)}$ fluctuate on a much shorter time scale.

The resulting contributions to the energy balance can be expressed in terms of the structure function tensor $D_{ik}(\mathbf{r})$. To simplify the expressions we introduce the second-order moment of the superscale velocity, $R_{ik}^{(\mathbf{r})}(\mathbf{r}')$, and the time-integrated correlation function of the subscale eddies, $N_{ik}^{(\mathbf{r})}(\mathbf{r}')$. Both can be expressed in terms of the structure function tensor:

$$\begin{split} R_{ik}^{(r)}(\mathbf{r}') &:= \langle \langle \! \langle u_i^{(r)}(\mathbf{x},t) u_k^{(r)}(\mathbf{x}+\mathbf{y}+\mathbf{r}',t) \rangle \! \rangle \rangle_y^{(r)} \\ &= \langle \! \langle u_i u_k \rangle \! \rangle - \frac{1}{2} \langle \langle \langle D_{ik}(\mathbf{r}'+\mathbf{y}_1+\mathbf{y}_2+\mathbf{y}_3) \rangle_{y_1}^{(r)} \rangle_{y_2}^{(r)} \rangle_{y_3}^{(r)}, \\ N_{ik}^{(r)}(\mathbf{r}') &:= \int_{-\infty}^t \mathrm{d}t' \langle \! \langle \tilde{u}_i^{(r)}(\mathbf{z},t) \tilde{u}_k^{(r)}(\mathbf{x}(t';\mathbf{z},t)+\mathbf{r}',t') \rangle \! \rangle. \end{split}$$

 $N_{ik}^{(r)}$ probes the (Lagrangian) dynamics and can be considered as an eddy transport coefficient for the superscale flow. To obtain a closed set of equations we express $N_{ik}^{(r)}$ in terms of equal time and therefore stationary *static* objects like the structure tensor $D_{ik}(\mathbf{r})$. This is achieved by continued fraction projector expansion (Grossmann & Thomae 1982; Daems *et al.* 1999). With the static subscale correlation $\tilde{C}_{ik}^{(r)}(\mathbf{x}') := \langle \tilde{u}_i^{(r)}(\mathbf{x},t)\tilde{u}_k^{(r)}(\mathbf{x}+\mathbf{x}',t) \rangle$ and the frequency matrix $\tilde{\Gamma}_{ik}^{(r)}(\mathbf{x}') := -\langle \langle \tilde{u}_i^{(r)}(\mathbf{z},t)d_t \tilde{u}_k^{(r)}(\mathbf{x}(t';\mathbf{z},t)+\mathbf{x}',t') \rangle |_{t'=t}$ we can write $N_{ik}^{(r)}$ in a 1-pole approximation as

$$N_{ik}^{(r)}(\mathbf{x}') = \tilde{C}_{ij}^{(r)}(\mathbf{x}')(\tilde{\Gamma}^{(r)}(\mathbf{x}'))_{jl}^{-1}\tilde{C}_{lk}^{(r)}(\mathbf{x}').$$

The tensor $\tilde{\mathbf{C}}$ and the frequency matrix $\tilde{\boldsymbol{\Gamma}}$ can be expressed in terms of the structure function tensor

$$\begin{split} \tilde{C}_{ik}^{(r)}(\mathbf{x}') &= -\frac{1}{2} \langle \langle D_{ik}(\mathbf{x}' + \mathbf{y}_1 + \mathbf{y}_2) \rangle_{y_1}^{(r)} \rangle_{y_2}^{(r)} + \langle D_{ik}(\mathbf{x}' + \mathbf{y}) \rangle_{y}^{(r)} - \frac{1}{2} D_{ik}(\mathbf{x}'), \\ \tilde{\Gamma}_{ik}^{(r)}(\mathbf{x}') &= \frac{2}{3} \epsilon \delta_{ik} - 2v \langle \Delta_y D_{ik}(\mathbf{y}) \rangle_{y}^{(r)} \\ &+ v \langle \langle \Delta_{y_1} D_{ik}(\mathbf{y}_1 + \mathbf{y}_2) \rangle_{y_1}^{(r)} \rangle_{y_2}^{(r)} + v \Delta_{x'} \tilde{C}_{ik}^{(r)}(\mathbf{x}'). \end{split}$$

For more details compare with the case of isotropic turbulence in Effinger & Grossmann (1987). In a general anisotropic case the dissipation matrix elements $v \langle \langle u_{i|j}^2 \rangle \rangle$ might be different for different *i*, *j*. Here, in the weakly anisotropic case, we assume that the anisotropy corrections are small on the scales where dissipation takes place. Therefore we insert one total dissipation rate per unit mass $\epsilon = v \langle \langle u_{i|j} u_{i|j} \rangle \rangle$ (summation implied). The superscale energy balance equation contains three contributions for the losses. $E_d(\mathbf{r})$ describes the direct viscous energy dissipation by the superscale eddies. The other two, consisting of a local and non-local part of $E_t(\mathbf{r})$, account for the energy transfer from the large to the small scales. These losses are balanced by the energy input rate $E_{in}(\mathbf{r})$ caused by the external forcing:

$$E_d(\mathbf{r}) + E_{t,lo}(\mathbf{r}) + E_{t,nolo}(\mathbf{r}) = E_{in}(\mathbf{r}).$$
(2.3)

As in Effinger & Grossmann (1987) the three different contributions can be written

$$E_d(\mathbf{r}) = \frac{1}{2} \nu \langle \langle \Delta_{y_1} D_{ii} (\mathbf{y}_1 + \mathbf{y}_2) \rangle_{y_1}^{(\mathbf{r})} \rangle_{y_2}^{(\mathbf{r})}, \qquad (2.4)$$

$$E_{t,lo}(\mathbf{r}) = -\frac{1}{2} N_{jk}^{(\mathbf{r})}(\mathbf{x}'=0) \partial_{x_j'} \partial_{x_k'} R_{ii}^{(\mathbf{r})}(\mathbf{x}')|_{\mathbf{x}'=0},$$
(2.5)

$$E_{t,nolo}(\mathbf{r}) = -\int d^{3}x' G(x') \partial_{x'_{i}} \partial_{x'_{j}} \{\partial_{x'_{j}} N_{lk}^{(\mathbf{r})}(\mathbf{x}') - \partial_{x'_{l}} N_{jk}^{(\mathbf{r})}(\mathbf{x}') \} \partial_{x'_{k}} R_{il}^{(\mathbf{r})}(\mathbf{x}') + \int d^{3}x' G(x') \Delta_{x'} (\partial_{x'_{i}} N_{lk}^{(\mathbf{r})}(\mathbf{x}')) \partial_{x'_{k}} R_{il}^{(\mathbf{r})}(\mathbf{x}').$$
(2.6)

The energy input rate is given by

$$E_{in}(\mathbf{r}) = \langle\!\langle u_i^{(\mathbf{r})} f_i^{(\mathbf{r})} \rangle\!\rangle. \tag{2.7}$$

Note that in contrast to the isotropic case all terms in the energy balance (2.3) now depend on the vector \mathbf{r} , not merely on its absolute value, the scale r.

Equation (2.3) together with (2.4)–(2.7) constitute a set of integro -differential equations for the tensor $D_{ik}(\mathbf{r})$. Now, anisotropy is assumed to be small. More precisely, in a SO(3)-decomposition of $D_{ik}(\mathbf{r})$ the *j*-amplitudes are assumed to decrease in magnitude for higher angular wavenumber *j*. Then (2.3) can be solved order by order in *j*, to give the structure function amplitudes $d_j(\mathbf{r})$. They will not be universal but depend on the anisotropy of the forcing. However, what we may hope is that the scaling of the individual *j*-amplitudes *is universal*. To analyse this, it is sufficient to focus on the scaling behaviour of the various contributions in (2.3).

Scalewise, multiple spatial averages can be reduced to first-order ones, e.g. $E_d(\mathbf{r}) \sim v/2\langle \Delta_y D_{ii}(\mathbf{y}) \rangle_y^{(\mathbf{r})}$, and local and non-local energy transfer rates scale with the same exponent. Here and in the following ~ has the meaning of 'scalewise equal'. Thus the energy balance equation scalewise simplifies to

$$E_{in}(\mathbf{r}) \sim \left\langle \frac{v}{2} \Delta_y D_{ii}(\mathbf{y}) \right\rangle_y^{(\mathbf{r})} + \left\langle \frac{\alpha}{\epsilon} D_{jl}(\mathbf{r}) D_{lk}(\mathbf{r}) \partial_{y_j} \partial_{y_k} D_{ii}(\mathbf{y}) \right\rangle_y^{(\mathbf{r})}.$$

Here, ϵ is the mean energy dissipation rate per unit mass, and the constant α takes into account the relative weight of the transport terms, E_t in (2.3). Scalewise this equation can be simplified even further:

$$E_{in}(\mathbf{r}) \sim \frac{1}{2} \left(v + \frac{\beta}{\epsilon} \mathbf{D}(\mathbf{r}) \mathbf{D}(\mathbf{r}) \right) \Delta \mathbf{D}(\mathbf{r}),$$
 (2.8)

where β takes into account the missing constants of proportionality. **DD** and ΔD stand for the tensorial products of two structure function tensors and of a second-order spatial derivative of the structure function tensor, respectively.

3. SO(3)-decomposition

Taking into account the full tensorial character of $D_{ik}(\mathbf{r})$ (equation (1.1)) complicates the resulting equation. Therefore, for simplicity, we assume that the *r*-scaling behaviour

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remains the same. As we are interested at present in the scaling exponents only, we disregard the tensorial character of the structure function (i.e. drop the index q of $d_{jmq}(r)$) and expand into spherical harmonics:

$$D(\mathbf{r}) \simeq d_{00}(r)Y_{00} + \sum_{m} d_{2m}(r)Y_{2m}(\hat{\mathbf{r}}) + \sum_{m} d_{4m}(r)Y_{4m}(\hat{\mathbf{r}}) + \dots$$

 $\sim \sum_{j} d_{j}(r) \sum_{m} Y_{jm}(\hat{\mathbf{r}}).$ (3.1)

Here, we assume that the scaling behaviour of $d_{jm}(r)$ is – for fixed j – the same for all m, and therefore simply write $d_j(r)$. We analogously expand the energy input rate into spherical harmonics:

$$E_{in}(\mathbf{r}) = \sum_{j,m} e_{jm}(r) Y_{jm}(\hat{\mathbf{r}}) \sim \sum_{j} e_{j}(r) \sum_{m} Y_{jm}(\hat{\mathbf{r}}), \qquad (3.2)$$

where

$$e_{jm}(\mathbf{r}) = \int \mathrm{d}(\cos\theta) \mathrm{d}\varphi \ Y_{jm}^*(\hat{\mathbf{r}}) E_{in}(\mathbf{r}).$$
(3.3)

Then we insert the SO(3)-decomposition (3.1) of the structure function and the corresponding expansion (3.2) of the energy input rate into equation (2.8).

From now on we only focus on the inertial subrange (ISR), $\eta \ll r \ll L$, where η is the Kolmogorov length, in which the second term on the right-hand side of equation (2.8) dominates. Thus the energy balance equation is

$$\sum_{j} e_{j}(r) \sum_{m} Y_{jm}(\hat{\boldsymbol{r}}) \sim \frac{\beta}{r^{2}} \left(\sum_{j} d_{j}(r) \sum_{m} Y_{jm}(\hat{\boldsymbol{r}}) \right)^{3}.$$
(3.4)

Projecting equation (3.4) on the different j-sectors and taking into account only the first three j (j = 0, 2, 4) yields three nonlinear equations for $d_0(r)$, $d_2(r)$ and $d_4(r)$:

$$e_{0}(r)r^{2} \sim \beta [A_{000,0}(d_{0}(r))^{3} + 3A_{022,0}d_{0}(r)(d_{2}(r))^{2} + 3A_{044,0}d_{0}(r)(d_{4}(r))^{2} + A_{222,0}(d_{2}(r))^{3} + 3A_{224,0}(d_{2}(r))^{2}d_{4}(r) + 3A_{244,0}d_{2}(r)(d_{4}(r))^{2} + A_{444,0}(d_{4}(r))^{3}], \qquad (3.5a)$$

$$e_{2}(r)r^{2} \sim \beta [3A_{002,2}(d_{0}(r))^{2}d_{2}(r) + 3A_{022,2}d_{0}(r)(d_{2}(r))^{2} + 3A_{044,2}d_{0}(r)(d_{4}(r))^{2} + 6A_{024,2}d_{0}(r)d_{2}(r)d_{4}(r) + A_{222,2}(d_{2}(r))^{3} + 3A_{224,2}(d_{2}(r))^{2}d_{4}(r) + 3A_{244,2}d_{2}(r)(d_{4}(r))^{2} + A_{444,2}(d_{4}(r))^{3}], \qquad (3.5b)$$

$$e_{4}(r)r^{2} \sim \beta [3A_{004,4}(d_{0}(r))^{2}d_{4}(r) + 3A_{022,4}d_{0}(r)(d_{2}(r))^{2} + 3A_{044,4}d_{0}(r)(d_{4}(r))^{2}$$

$$e_4(r)r^2 \sim \beta [3A_{004,4}(d_0(r))^2 d_4(r) + 3A_{022,4}d_0(r)(d_2(r))^2 + 3A_{044,4}d_0(r)(d_4(r))^2 + 6A_{024,4}d_0(r)d_2(r)d_4(r) + A_{222,4}(d_2(r))^3 + 3A_{224,4}(d_2(r))^2 d_4(r) + 3A_{244,4}d_2(r)(d_4(r))^2 + A_{444,2}(d_4(r))^3].$$
(3.5c)

Here, $A_{j_1 j_2 j_3, j_4} = \sum_{m_1, m_2, m_3, m_4} \int d(\cos \theta) d\varphi Y_{j_4 m_4}^* Y_{j_1 m_1} Y_{j_2 m_2} Y_{j_3 m_3}$. The $A_{j_1 j_2 j_3, j_4}$ can have either sign.

To extract the scaling laws for the different $d_j(r)$, equations (3.5) have to be solved. But before doing so, we have to specify the energy input rate $E_{in}(\mathbf{r})$, equation (2.7), which depends on the external forcing $f_i^{(r)}$.

4. Anisotropic forcing

In the isotropic and homogeneous case $E_{in}(\mathbf{r}) = E_{in}$ is a scale-independent constant (Effinger & Grossmann (1987)). The reason is the following. While the superscale



FIGURE 1. Scaling behaviour of the amplitudes of the second-order structure function $d_0(r)$, $d_2(r)$ and $d_4(r)$, for an analytic forcing. (a) Strong isotropic forcing together with weak anisotropy corrections, $e_0/\epsilon = 0.89$, $e_2/\epsilon = 0.1$, $e_4/\epsilon = 0.01$. (b) Strong anisotropic forcing: the first anisotropic sector j = 2 dominates the energy input, $e_0/\epsilon = 0.001$, $e_2/\epsilon = 0.989$, $e_4/\epsilon = 0.01$. This might already reach the limits of our assumptions regarding weak anisotropy. The dip of the d_0 -curve originates from a change of sign of $d_0(r)$.

velocity field $u_i^{(r)}$ contains all scales larger than r, the forcing $f_i^{(r)}$ has the outer scale L only. For each $r \leq L$ the complete forcing is included in the same and therefore r-independent way. Of course, $E_{in} = \epsilon$. In the present case, however, the forcing has to provide an anisotropic flow. As a consequence we shall find that $f_i^{(r)}$ has to depend on all scales r, implying that the energy input rate will also depend on all r.

We will discuss two different classes of anisotropic flows: a general analytic forcing and a non-analytic forcing. For both we can determine the scaling behaviour with dimensional arguments.

4.1. Analytic forcing

Let us assume that the forcing $f_i(\mathbf{x}) \sim a_i \mathbf{k} \cdot \mathbf{r} \sin(\mathbf{k} \cdot \mathbf{x})$ and the velocity profile $u_i(\mathbf{x}) \sim b_i \mathbf{k} \cdot \mathbf{r} \sin(\mathbf{k} \cdot \mathbf{x})$ depend on one input wavenumber \mathbf{k} only. They are analytic in the components of position \mathbf{x} and the scale vector \mathbf{r} . To fulfil the incompressibility condition, $\partial_i f_i = 0$ and $\partial_i u_i = 0$, the vectors a_i and b_i must be chosen as $a_i k_i = b_i k_i = 0$. Then, applying the y-average defined in equation (2.1) yields $u_i^{(r)} \sim f_i^{(r)} \sim [\cos(\mathbf{k} \cdot (\mathbf{x} + \mathbf{r})) - \cos(\mathbf{k} \cdot (\mathbf{x} - \mathbf{r}))]$. Therefore

$$E_{in}(\mathbf{r}) \sim \left\langle \left\langle \left(\cos\left(\mathbf{k} \cdot (\mathbf{x} + \mathbf{r})\right) - \cos\left(\mathbf{k} \cdot (\mathbf{x} - \mathbf{r})\right)\right)^2 \right\rangle \right\rangle = 1 - \cos\left(2kr\xi\right)$$
(4.1)

with $\xi = \cos \theta$, the projection on the z-axis defined by \hat{k} . A power series expansion of $E_{in}(\mathbf{r})$ in the variable $r\xi$ inserted into equation (3.3) implies (because $\xi^n \perp Y_{jm}$ for all n < j) that $e_{jm} \sim r^j$ plus higher powers.

We now solve equations (3.5) and extract the power laws for the different $d_j(r)$. Figure 1 shows the solutions of (3.5). In (a) the isotropic part of the energy input e_0 is the largest one, and the anisotropy contributions are small corrections. In this case, over the whole calculated range $10^{-4} \leq r/L \leq 1$ the $d_j(r)$ scale as

$$d_i(r) \sim r^{j+2/3}$$
. (4.2)

We can see this scaling behaviour easily from equations (3.5): since $d_4 \ll d_2 \ll d_0$, the dominating term on the right hand side of equation (3.5*a*) is $A_{000,0}(d_0)^3$. It is balanced by e_0r^2 . Therefore, $d_0 \sim r^{2/3}$. Then, in equations (3.5*b*, *c*) the leading terms $A_{002,2}(d_0)^2 d_2$ and $A_{004,4}(d_0)^2 d_4$ are balanced by $e_2r^2 \sim r^4$ and $e_4r^2 \sim r^6$, respectively. Therefore, we expect $d_2 \sim r^{8/3}$ and $d_4 \sim r^{14/3}$. Though for j = 0 we recover the mean-field scaling of the isotropic amplitude of the structure function $d_0 \sim r^{2/3}$, as in Effinger & Grossmann (1987), the result for the j = 2 sector is at variance with the experimental finding by Kurien *et al.* (2000), who found a scaling exponent close to 4/3. If, on the other hand, we chose a strongly anisotropic energy input with $e_2 \gg e_0, e_4$ (in which case the assumption of weak anisotropy of course breaks down), then the *r*-scaling range with $d_j(r) \sim r^{j+2/3}$ becomes smaller, while at larger values of *r* a new scaling range $d_j \sim r^{4/3}$ with the same exponent 4/3 for all *j* emerges, see figure 1(*b*). For j = 2 this finding is now consistent with the experimental observations by Kurien *et al.* (2000). However, it is inconsistent with the exponent 2/3 to be expected for the j = 0 amplitude. In summary, the analytic energy input does not seem to describe the experimental findings. We therefore now explore the option of non-analytic forcing.

4.2. Non-analytic forcing

We consider a shear flow with its shear in the 3-direction. Then the three *f*-components are different. We decompose the velocity u_i and the forcing f_i into an isotropic (*iso*) and a (smaller) anisotropic (*an*) part: $u_i = u_i^{(iso)} + u_i^{(an)}$, $f_i = f_i^{(iso)} + f_i^{(an)}$. Then, at first order of anisotropy

$$\langle \langle u_i^{(r)} f_i^{(r)} \rangle \rangle \simeq \langle \langle u_i^{(r)} f_i^{(r)(iso)} \rangle \rangle + \langle \langle u_i^{(r)(iso)} f_i^{(r)(an)} \rangle \rangle = E_{in}^{(iso)} + E_{in}^{(an)}(r).$$
(4.3)

Repeating the arguments at the beginning of §4 for the isotropic case the first term on the right-hand side does not depend on r, i.e. $E_{in}^{(iso)} \sim r^0$. Namely, since $f_i^{(r)(iso)}$ has scales of order L only, the smaller scales in the products with $u_i^{(r)(iso)}$ or $u_i^{(r)(an)}$ cannot contribute, irrespective of their degree of isotropy. The second term, however, will depend on r and introduces anisotropy.

Let us determine $E_{in}^{(an)}(\mathbf{r})$ by scaling arguments. The flow profile in shear flow is generated by the boundary conditions: one plate is moving with velocity U, the other one is at rest. These boundary conditions have to be mimicked by the forcing f in an infinitely extending flow. The linear mean velocity profile (U/L)z (and therefore also the corresponding f) has Fourier coefficients on all scales,

$$u^{(an)}(k) = \frac{U}{L^2} \int_{-L}^{L} dz \ z e^{ikz} = 2iU\left(\frac{\sin kL}{k^2L^2} - \frac{\cos kL}{kL}\right).$$

In the case of large k, i.e. $k^{-1} \sim z \ll L$, the second term dominates. We therefore asymptotically find

$$u^{(an)}(k) \sim \frac{\cos kL}{kL} \sim \frac{1}{k} \sim z = r\cos\theta.$$
(4.4)

Incidentally, a parabolic velocity profile as in pipe flow, $(U/L^2)z^2$, gives the same asymptotic scaling, $u^{(an)}(k) \sim (\sin kL)/kL \sim 1/k \sim z$ for large k.

Next, we determine the *r*-dependence of $f^{(an)}$. From the Navier–Stokes equation we have $\partial u/\partial t = \cdots + f$. Therefore, the dimension and *r*-scaling of *f* must correspond to that of u/τ , where τ is the *r*-eddy turnover time. In the isotropic case the turnover time τ scales like $\tau(r) \sim r/u(r) \sim r/r^{1/3} \sim r^{2/3}$. We use the *r*-dependence of the anisotropic velocity field $u^{(an)}(r)$ together with that of the isotropic turnover time $\tau(r)$ to estimate the scaling of the anisotropic forcing $f^{(an)}$ in first order. Since $u^{(an)}$ behaves as $u^{(an)} \sim r \cos \theta$ according to equation (4.4), we have $f^{(an)} \sim r(\cos \theta)/r^{2/3} \sim r^{1/3} \cos \theta$. Note that this anisotropic forcing scales as the isotropic velocity $u^{(iso)} \sim r^{1/3}$. Then



FIGURE 2. Scale dependence of the amplitudes $d_0(r)$, $d_2(r)$ and $d_4(r)$ of the second-order structure function, for a non-analytic forcing. The forcing is assumed to be predominantly isotropic with small anisotropy corrections, $e_0/\epsilon = 0.89$, $e_2/\epsilon = 0.1$, $e_4/\epsilon = 0.01$.

both factors of $E_{in}^{(an)}(\mathbf{r})$ in (4.3) not only contain all scales, but also the same power-law exponents. We therefore find

$$E_{in}^{(an)}(\mathbf{r}) \sim r^{2/3} \cos \theta. \tag{4.5}$$

The forcing still has an additional factor $g(\theta, \phi)$, which can be chosen such that the incompressibility constraint $\partial_i f_i = 0$ is fulfilled. Knowing now the scaling behaviour of the energy input rate, we proceed to expand it into spherical harmonics (see equation (3.2)),

$$E_{in}(\mathbf{r}) = E_{in}^{(iso)} + r^{2/3} \tilde{E}_{in}^{(an)}(\hat{\mathbf{r}}) = \sum_{j,m} e_{jm}(r) Y_{jm}(\hat{\mathbf{r}}).$$

Here, $e_{00}(r)$ equals $E_{in}^{(iso)}$, while for j > 0 the input rate amplitudes are $e_{jm}(r) = \int d(\cos\theta)d\phi \ Y_{jm}^*(\hat{r})E_{in}^{(an)}(r)$. Here the *r*- and ξ -dependences are not coupled as $r\xi$, in contrast to the analytic case treated above. Thus there is no *j*-dependence of the leading *r*-power of the input amplitudes e_{jm} . The lowest *j*-value projection $e_{jm}(r)$ of the anisotropy correction, in general j = 2, and all the higher ones, have the same power law, here according to $(4.5) \sim r^{2/3}$. The physics behind this decoupling of the *r*- and ξ -dependence is that a shear profile – in contrast to a single input wavenumber k – contains all wavenumbers. The *r*-dependence is even non-analytical. Hence, no expansion in $r\xi$ with only integer powers holds.

From equations (3.5) and with $d_4 \ll d_2 \ll d_0$, we find that the leading terms in equations (3.5*b*,*c*) are $(d_0)^2 d_2 \sim r^{8/3}$ and $(d_0)^2 d_4 \sim r^{8/3}$. With $d_0 \sim r^{2/3}$ from equation (3.5*a*) this leads to $d_2 \sim d_4 \sim r^{4/3}$. The solutions of equations (3.5) for this non-analytic forcing describing shear flow are shown in figure 2. For j = 0 we recover the isotropic scaling of the structure function $d_0(r) \sim r^{2/3}$. However, for all higher amplitudes we obtain

$$d_i(r) \sim r^{4/3}$$
 (4.6)

in a wide range of r. We have shown only the case where the isotropic forcing dominates, $e_0 \gg e_2 \gg e_4$. If the anisotropic contributions to the energy input increase, the r-range of scaling behaviour $d_0 \sim r^{2/3}$ and $d_j \sim r^{4/3}$ for all $j \ge 2$ is again shifted towards smaller r as in the case of analytic forcing.

As was argued by L'vov & Procaccia (1995*a*, *b*, 1996), in the exact resummation theory of Navier–Stokes turbulence no infrared (IR) divergence occurs if all *j*-factor scaling exponents ζ^{j} are bounded by 4/3. Otherwise IR-divergences cannot

be excluded. Our results in the non-analytic case therefore exclude IR-divergences while in the previous analytic case they cannot be ruled out. Another possibility for controlling these IR divergences, in spite of second-order moment exponents being larger than 4/3, is by limiting the forcing to scales smaller than L, as recently shown in Kraichnan-model-type dynamics for a vector field, see Procaccia & Arad (2001).

5. Summary

Within a variable-scale mean-field theory of the Navier–Stokes equation we have derived the scaling exponents of the different *j*-amplitudes of the SO(3)-decomposition of the second-order structure function for weakly anisotropic turbulent flow. The limitation of this approach is its mean-field character. Clearly, intermittency effects cannot be captured, but we consider those to be small for second-order moments, to which the method is limited anyway. In the isotropic sector j = 0 we recover the classical scaling behaviour $\sim r^{2/3}$. The higher-order contributions, i.e. the anisotropic parts of the flow field, can be calculated order by order in Y_{jm} . They yield, for all *j*, the same mean-field scaling $r^{4/3}$ for a non-analytic forcing might be more general, and therefore valid for a larger variety of anisotropic flows. Moreover, only the results for the non-analytic forcing are consistent with existing experimental measurements for the j = 0 and j = 2 amplitudes.

The work is part of the research program of FOM, which is financially supported by NWO. It was also supported by the German-Israeli Foundation (GIF) and by the European Union (EU) under contract HPRN-CT-2000-00162.

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